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Doug Ravenel University of Rochester

March 22, 2024

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March 7, 1970

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A solar eclipse occurs when the shadow of the moon falls on the earth.

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A solar eclipse occurs when the shadow of the moon falls on the earth. It can be up to 250 miles wide and moves eastward (relative to the surface of the eastwardly rotating earth) at varying speeds over 1000 MPH.

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A lunar eclipse occurs when the moon passes through the earth's shadow, olar eclipses: Why they happen when they do



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A lunar eclipse occurs when the moon passes through the earth's shadow, moving at roughly 2300 MPH. Solar eclipses: Why they happen when they do



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The eclipse of March 8-9, 2016 seen from space. The moon's shadow is over Indonesia. olar eclipses: Why they happen when they do



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The diameter of the moon is 2,159 miles (3,474 km).

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The Moon is on average of 238,855 miles (384,400 km) away from Earth.

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The Moon is on average of 238,855 miles (384,400 km) away from Earth. The average distance between Earth and the Sun is about 93 million miles (150 million km) about 400 times that to the moon.

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The Moon is on average of 238,855 miles (384,400 km) away from Earth. The average distance between Earth and the Sun is about 93 million miles (150 million km) about 400 times that to the moon.

Why are these two ratios the same?

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Why are these two ratios the same? Lucky coincidence!

No other moon in the solar system has this property.

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The sun is both 400 times as wide and 400 times as far away as the moon.

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The sun is both 400 times as wide and 400 times as far away as the moon. This means they appear to be the same size in the sky. Solar eclipses: Why they happen when they do



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The sun is both 400 times as wide and 400 times as far away as the moon. This means they appear to be the same size in the sky.

The diameter of each is about .9 percent of its distance.

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The sun is both 400 times as wide and 400 times as far away as the moon. This means they appear to be the same size in the sky.

The diameter of each is about .9 percent of its distance. That makes its angular diameter about half a degree, or 30 minutes of arc, written as 30'.

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This means that when the moon passes directly in front of the sun, it is just big enough to cover it,

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The diameter of each is about .9 percent of its distance. That makes its angular diameter about half a degree, or 30 minutes of arc, written as 30'.

This means that when the moon passes directly in front of the sun, it is just big enough to cover it, but not big enough to cover the snowy corona around it.





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The distance to the moon varies from 225,700 miles (perigee) to 251,900 miles (apogee).

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The distance to the moon varies from 225,700 miles (perigee) to 251,900 miles (apogee). This means its angular diameter varies by about 10 percent. If it passes in front of the sun when it is too far away,

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The distance to the moon varies from 225,700 miles (perigee) to 251,900 miles (apogee). This means its angular diameter varies by about 10 percent. If it passes in front of the sun when it is too far away, it is not big enough to cover the sun completely.

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This results in an annular eclipse.

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This results in an annular eclipse. The sun is still shining and too bright to look at even when the moon is dead center. You could miss it completely! Solar eclipses: Why they happen when they do



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There was one in Rochester on May 10, 1994.

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There was one in Rochester on May 10, 1994. We are lucky to live in a place with two solar eclipses in just 30 years! Solar eclipses: Why they happen when they do



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This results in an annular eclipse. The sun is still shining and too bright to look at even when the moon is dead center. You could miss it completely!



There was one in Rochester on May 10, 1994. We are lucky to live in a place with two solar eclipses in just 30 years! The image to the left was photographed through a filter. Solar eclipses: Why they happen when they do



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On the other hand, if the moon is closer than average, it will appear to be slightly larger than the sun.

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On the other hand, if the moon is closer than average, it will appear to be slightly larger than the sun. This means that totality will last a little longer. Solar eclipses: Why they happen when they do



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On the other hand, if the moon is closer than average, it will appear to be slightly larger than the sun. This means that totality will last a little longer. This will be the case on April 8. Solar eclipses: Why they happen when they do



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On the other hand, if the moon is closer than average, it will appear to be slightly larger than the sun. This means that totality will last a little longer. This will be the case on April 8.



https://www.timeanddate.com/eclipse/solar/2024-april-8

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WTF fun fact #10698

If you were on the moon, the Earth wouldn't actually move in the sky. It would appear to wobble a little because the moon is elliptical but it would never "rise" or "set".





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The earth would always be in the same spot in the sky.

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The earth would appear to be four times larger than the moon appears to be from earth,

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The earth would appear to be four times larger than the moon appears to be from earth, because it is!

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The earth would appear to be four times larger than the moon appears to be from earth, because it is! Its angular diameter would be 2 degrees instead of the moon and sun's half a degree as seen from earth. Solar eclipses: Why they happen when they do



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The earth would appear to be four times larger than the moon appears to be from earth, because it is! Its angular diameter would be 2 degrees instead of the moon and sun's half a degree as seen from earth. The sun's angular diameter would be the same as it is on earth. Solar eclipses: Why they happen when they do



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The earth would never move but you could see it rotate once every 24 hours.

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The earth would never move but you could see it rotate once every 24 hours. (Seen from earth, the moon moves but does rotate.) Solar eclipses: Why they happen when they do



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The earth would appear to be four times larger than the moon appears to be from earth, because it is! Its angular diameter would be 2 degrees instead of the moon and sun's half a degree as seen from earth. The sun's angular diameter would be the same as it is on earth.

The earth would never move but you could see it rotate once every 24 hours. (Seen from earth, the moon moves but does rotate.) The earth would have phases (new earth, full earth, crescent earth, etc) just like the moon does on earth. Solar eclipses: Why they happen when they do



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The earth would always be in the same spot in the sky. This is because the rotation of the moon is locked in by gravity to its orbital motion about the earth. The same side of it is always facing us. This is not a coincidence.

The earth would appear to be four times larger than the moon appears to be from earth, because it is! Its angular diameter would be 2 degrees instead of the moon and sun's half a degree as seen from earth. The sun's angular diameter would be the same as it is on earth.

The earth would never move but you could see it rotate once every 24 hours. (Seen from earth, the moon moves but does rotate.) The earth would have phases (new earth, full earth, crescent earth, etc) just like the moon does on earth. The sun would rise and set once a month. Solar eclipses: Why they happen when they do



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Sometimes (when there is a lunar eclipse on earth) the sun would be eclipsed by the earth.

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(up to 4 hours) than a solar eclipse on earth,

Sometimes (when there is a lunar eclipse on earth) the sun would be eclipsed by the earth. Totality would last much longer Solar eclipses: Why they happen when they do



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Sometimes (when there is a lunar eclipse on earth) the sun would be eclipsed by the earth. Totality would last much longer (up to 4 hours) than a solar eclipse on earth, but you would only see a small portion of the corona

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Sometimes (when there is a lunar eclipse on earth) the sun would be eclipsed by the earth. Totality would last much longer (up to 4 hours) than a solar eclipse on earth, but you would only see a small portion of the corona and only at the start and end of totality.

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https://en.wikipedia.org/wiki/Solar_eclipses_on_the_Moon

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Back to eclipses on earth. The plane of the moon's orbit is inclined from that of the earth's orbit by 5.145 degrees.

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5° tilt of moon's

orbital plane

Back to eclipses on earth. The plane of the moon's orbit is inclined from that of the earth's orbit by 5.145 degrees. Usually the moon passes north or south of the sun at new moon, and north or south of earth's shadow at full moon.

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Back to eclipses on earth. The plane of the moon's orbit is inclined from that of the earth's orbit by 5.145 degrees. Usually the moon passes north or south of the sun at new moon, and north or south of earth's shadow at full moon. The points where the moon crosses the earth's orbital plane are the ascending node Ω and descending node \Im .

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Each red line in the picture is the line of nodes. Eclipses can occur only during the time of year when it is pointing toward the sun (within a few degrees), known as eclipse season.

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Each red line in the picture is the line of nodes. Eclipses can occur only during the time of year when it is pointing toward the sun (within a few degrees), known as eclipse season. It occurs roughly twice each calendar year.

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Meanwhile the line of nodes itself rotates clockwise roughly every 18.60 years.

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Over 2000 years ago Babylonian astronomers noticed that eclipses (both lunar and solar) occur in cycles of roughly 18 years.

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A solar/lunar eclipse is almost always followed by a lunar/solar eclipse half a saros or sar (roughly 9 years) later.

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March 7, 1970

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April 8, 2024 March 7, 1970 A nice thing about this Saros series is that the eclipses in it are getting longer.

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April 8, 2024 March 7, 1970 A nice thing about this Saros series is that the eclipses in it are getting longer. The one next month will last longer than the one I saw 54 years ago.

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In 1991 there was a solar eclipse visible in Hawaii and Mexico.

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July 11, 1991

In 1991 there was a solar eclipse visible in Hawaii and Mexico. The Keck observatory atop Mauna Kea on the big island was in its path, and skies were clear there olar eclipses: Why they happen when they do



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July 11, 1991

In 1991 there was a solar eclipse visible in Hawaii and Mexico. The Keck observatory atop Mauna Kea on the big island was in its path, and skies were clear there but some beaches on the island were clouded out.

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July 11, 1991

There was an algebraic topology conference in Mexico organized around the eclipse at a site south of Mexico City.

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There was an algebraic topology conference in Mexico organized around the eclipse at a site south of Mexico City. We had clear skies, after several days of clouds.

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July 11, 1991

Teotihuacan is a spectacular archeaological site north of Mexico City.

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July 11, 1991

Teotihuacan is a spectacular archeaological site north of Mexico City. Ten thousand people gathered there to watch the eclipse.

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Those ten thousand tourists were not only clouded out,

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Those ten thousand tourists were not only clouded out, the eclipse made it rain on them!

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Those ten thousand tourists were not only clouded out, the eclipse made it rain on them! Totality causes a small temperature drop.

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July 11, 1991 May 29, 1919 Four saros periods back from the 1991 eclipse gets us to May 29, 1919,

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July 11, 1991 May 29, 1919 Four saros periods back from the 1991 eclipse gets us to May 29, 1919, the day of the Eddington-Einstein eclipse,

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July 11, 1991 May 29, 1919 Four saros periods back from the 1991 eclipse gets us to May 29, 1919, the day of the Eddington-Einstein eclipse, possibly the most famous one in the history of science.

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In his 1915 paper on general relativity, Einstein predicted that light from a distant star passing very near the sun would be bent by the sun's gravity. olar eclipses: Why they happen when they do



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In his 1915 paper on general relativity, Einstein predicted that light from a distant star passing very near the sun would be bent by the sun's gravity. This change in direction would be very small, only 1.75 seconds or arc, or .000486 degrees.

Sun

Apparent position

Actual

position of star

Earth

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In his 1915 paper on general relativity, Einstein predicted that light from a distant star passing very near the sun would be bent by the sun's gravity. This change in direction would be very small, only 1.75 seconds or arc, or .000486 degrees. It would lead to an apparent change in the star's position.

Sun

Apparent position

Actual

position of star



In his 1915 paper on general relativity, Einstein predicted that light from a distant star passing very near the sun would be bent by the sun's gravity. This change in direction would be very small, only 1.75 seconds or arc, or .000486 degrees. It would lead to an apparent change in the star's position. Because of the sun's brightness, this change could only be observed during a total solar eclipse.

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Arthur Eddington 1882-1944 Albert Einstein 1879-1955 Photo by Eddington

The British physicist Arthur Eddington used the 1919 eclipse to make such an observation from the island of Principé off the west coast of Africa.

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Arthur Eddington 1882-1944 Albert Einstein 1879-1955 Photo by Eddington

The British physicist Arthur Eddington used the 1919 eclipse to make such an observation from the island of Principé off the west coast of Africa. Just six months after the end of World War I, a British scientist travelled to the ends of the earth to confirm the theory of a German one.

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Arthur Eddington 1882-1944

Albert Einstein 1879-1955

Photo by Eddington

The British physicist Arthur Eddington used the 1919 eclipse to make such an observation from the island of Principé off the west coast of Africa. Just six months after the end of World War I, a British scientist travelled to the ends of the earth to confirm the theory of a German one. The measurement was as Einstein had predicted.

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Arthur Eddington 1882-1944

Albert Einstein 1879-1955

Photo by Eddington

The British physicist Arthur Eddington used the 1919 eclipse to make such an observation from the island of Principé off the west coast of Africa. Just six months after the end of World War I, a British scientist travelled to the ends of the earth to confirm the theory of a German one. The measurement was as Einstein had predicted. The ensuing headlines made him the most famous scientist in the world.



August 11, 1999

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August 11, 1999

In 1999 there was an eclipse visible across Europe.

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August 11, 1999

In 1999 there was an eclipse visible across Europe. I was hoping for a conference organized around it like the one in 1991.

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August 11, 1999

In 1999 there was an eclipse visible across Europe. I was hoping for a conference organized around it like the one in 1991. It did not happen, which was just as well. olar eclipses: Why they happen when they do



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August 11, 1999

In 1999 there was an eclipse visible across Europe. I was hoping for a conference organized around it like the one in 1991. It did not happen, which was just as well. Most of western Europe was cloudy that day.

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Most of western Europe was cloudy that day.





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Most of western Europe was cloudy that day. One exception was Bucharest, where the skies were clear.



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August 11, 1999



August 21, 2017

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The next eclipse in the series was visible across the United States from Oregon to South Carolina.



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The next eclipse in the series was visible across the United States from Oregon to South Carolina. There was an another algebraic topology conference organized around it in Portland, and skies were clear.



Courtesy of "Totality - Eclipses of the Sun" by Littmann, Willcox and Espenak

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To understand the Saros cycle we need to know more about how the moon moves.

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1. A sidereal month (27.32166 days or 27 days, 7 hours, 43 minutes, 11.6 seconds) is the time the moon takes to make one complete turn around the earth.

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1. A sidereal month (27.32166 days or 27 days, 7 hours, 43 minutes, 11.6 seconds) is the time the moon takes to make one complete turn around the earth.

2. A synodic month (29.53059 days or 29 days, 12 hours, 44 minutes, 2.8 seconds) is the interval between two new (or full) moons.

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One sidereal month after a new moon, the moon is not yet new again

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One sidereal month after a new moon, the moon is not yet new again because the earth has advanced in its orbit around the sun.

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One sidereal month after a new moon, the moon is not yet new again because the earth has advanced in its orbit around the sun. It needs 2 more days to reach the new earthsun line.

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1. A sidereal month (27.32166 days) is the time the moon goes around the earth once in the stellar frame of reference.

3. A draconic month (27.21222 days) is the time between ascending (or descending) nodes,

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3. A draconic month (27.21222 days) is the time between ascending (or descending) nodes, the point where the moon crosses the earth's orbital plane moving north (or south).

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The draconic month is roughly 2.6 hours shorter than the sidereal month

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The draconic month is roughly 2.6 hours shorter than the sidereal month because the line of nodes is rotating slowly westward (clockwise), making a complete turn every 18.60 years.

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The draconic month is roughly 2.6 hours shorter than the sidereal month because the line perpendicular to the plane of the moon's orbit (and with it the line of nodes) is rotating clockwise roughly every 18.6 years.

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1. A sidereal month (27.32166 days) is the time the moon goes around the earth once in the stellar frame of reference.

4. An anomalistic month (27.55455 days) is the interval between two perigees, the times when the moon is closest to the earth.

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1. A sidereal month (27.32166 days) is the time the moon goes around the earth once in the stellar frame of reference.

4. An anomalistic month (27.55455 days) is the interval between two perigees, the times when the moon is closest to the earth.

The anomalistic month is roughly 5.6 hours longer than the sidereal month because the perigee/apogee line,

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1. A sidereal month (27.32166 days) is the time the moon goes around the earth once in the stellar frame of reference.

4. An anomalistic month (27.55455 days) is the interval between two perigees, the times when the moon is closest to the earth.

The anomalistic month is roughly 5.6 hours longer than the sidereal month because the perigee/apogee line, called the line of apsides,

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The rotation is called apsidal precession.

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1. A sidereal month (27.32166 days) is the time the moon goes around the earth once in the stellar frame of reference.

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2. A synodical month (29.53059 days) is the interval between two new (or full) moons.

3. A draconic month (27.21222 days) is the time between ascending (descending) nodes, the point where the moon crosses the earth's orbital plane moving north (south).

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The saros period, 6,585.321347 days, which is 18 years plus 10, 11 or 12 calendar days,

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The saros period, 6,585.321347 days, which is 18 years plus

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which is 241.999 (called this number δ) draconic months

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The saros period, 6,585.321347 days, which is 18 years plus 10, 11 or 12 calendar days, is precisely 223 synodic months, which is 241.999 (called this number δ) draconic months and 238.992 (call this number α) anomalistic months.

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Why 223? It is the smallest number of synodic months that is this close to whole multiples of both the draconic and anomalistic months.

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In a series of lunar eclipses where the moon is near \Im , the first one (which is only partial) occurs when the moon as passed \Im and crosses the southern edge of the earth's shadow.

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In a series of lunar eclipses where the moon is near \Im , the first one (which is only partial) occurs when the moon as passed \Im and crosses the southern edge of the earth's shadow. The Moon's path is shifted slighly northward after each successive saros.

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The northward movement occurs because each saros is slightly less (241.999) than a whole number of draconic months.

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A typical Saros series has 71 eclipses and lasts 1262 years.

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Sidebar: Living on the moon

Why isn't there an eclipse every month?

The Saros series

Examples of Saros series

Example 1: Saros series 139

Example 2: Saros series 136

Example 3: Saros series 145

Why the Saros cycle?

Future eclipses

The planets on April 8

THANK YOU!

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